



## 5.4 Periodic noise reduction by frequency domain filtering

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### 5.4.1 Bandreject Filters

#### Ideal bandreject filter

$$H(u, v) = \begin{cases} 1, & D(u, v) < D_0 - \frac{W}{2} \\ 0, & D_0 - \frac{W}{2} \leq D(u, v) \leq D_0 + \frac{W}{2} \\ 1, & D(u, v) > D_0 + \frac{W}{2} \end{cases}$$

$W$  is the width of the band

#### Butterworth bandreject filter

$$H(u, v) = \frac{1}{1 + \left\{ \frac{D(u, v) W}{D^2(u, v) - D_0^2} \right\}^{2n}}$$

#### Gaussian bandreject filter

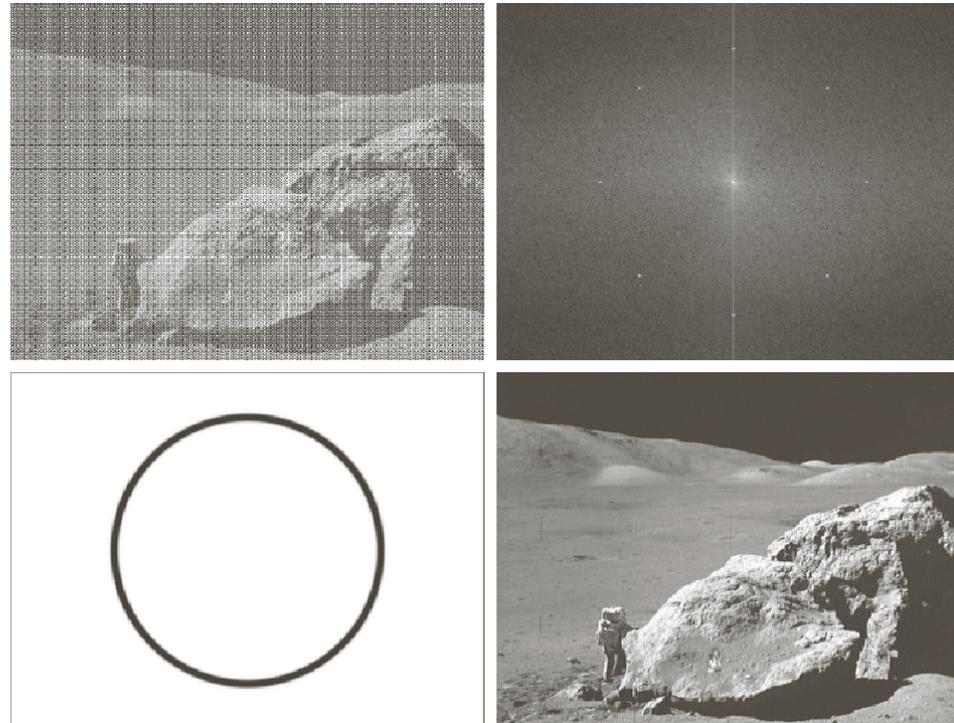
$$H(u, v) = 1 - e^{-\left\{ \frac{D^2(u, v) - D_0^2}{D(u, v) W} \right\}}$$



**FIGURE 5.15** From left to right, perspective plots of ideal, Butterworth (of order 1), and Gaussian bandreject filters.

a b  
c d

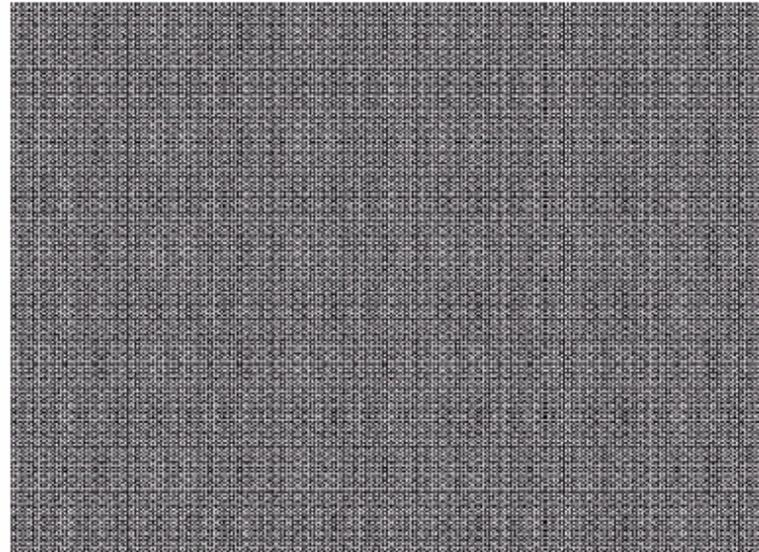
**FIGURE 5.16**  
(a) Image corrupted by sinusoidal noise.  
(b) Spectrum of (a).  
(c) Butterworth bandreject filter (white represents 1).  
(d) Result of filtering.  
(Original image courtesy of NASA.)



## 5.4.2 Bandpass Filters

$$H_{BP}(u, v) = 1 - H_{BR}(u, v)$$

**FIGURE 5.17**  
Noise pattern of  
the image in  
Fig. 5.16(a)  
obtained by  
bandpass filtering.



## 5.4.3 Notch Filters

### Ideal notch reject filter

$$H(u, v) = \begin{cases} 0, & D_1(u, v) \leq D_0 \text{ or } D_2(u, v) \leq D_0 \\ 1, & \text{otherwise} \end{cases}$$

$$D_1(u, v) = [ (u - M/2 - u_0)^2 + (v - N/2 - v_0)^2 ]^{1/2}$$

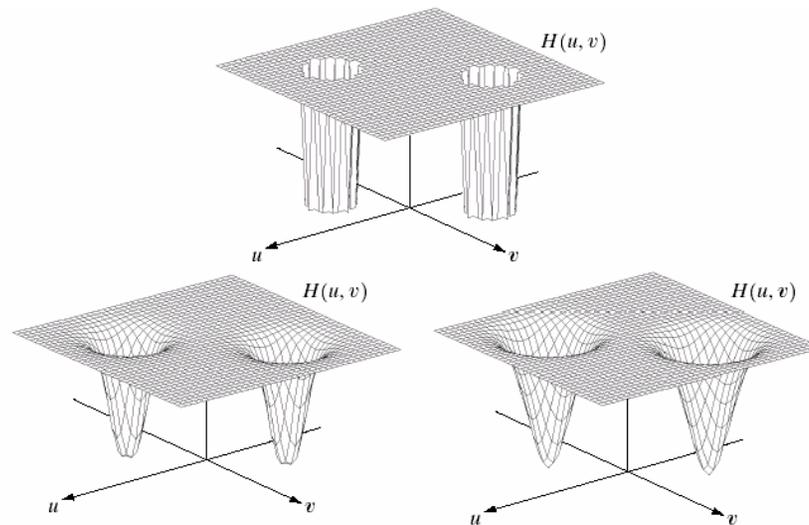
$$D_2(u, v) = [ (u - M/2 + u_0)^2 + (v - N/2 + v_0)^2 ]^{1/2}$$

## Butterworth notch reject filter

$$H(u, v) = \frac{1}{1 + \left\{ \frac{D_0^2}{D_1(u, v) D_2(u, v)} \right\}^{2n}}$$

## Gaussian notch reject filter

$$H(u, v) = 1 - e^{-\left\{ \frac{D_1(u, v) D_2(u, v)}{D_0^2} \right\}}$$



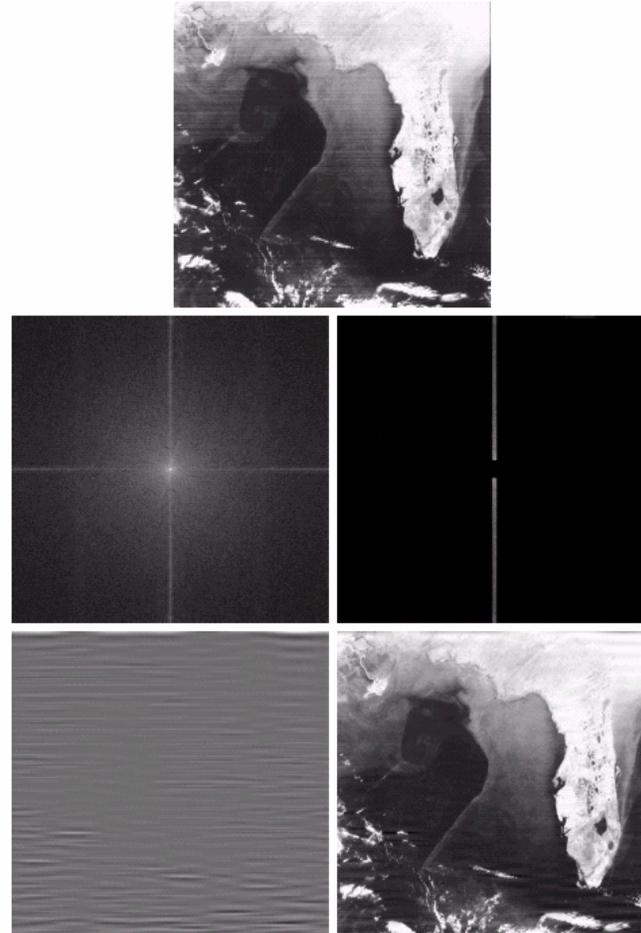
a  
b c

**FIGURE 5.18** Perspective plots of (a) ideal, (b) Butterworth (of order 2), and (c) Gaussian notch (reject) filters.

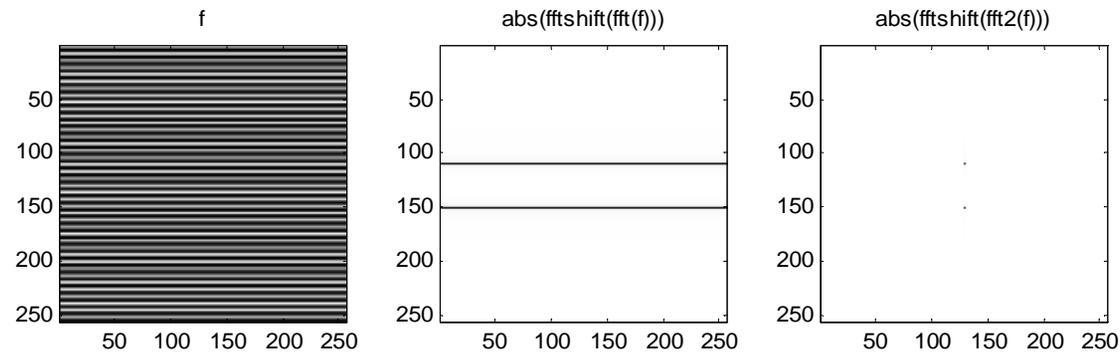
## Notch pass filters

$$H_{\text{NP}}(u, v) = 1 - H_{\text{NR}}(u, v)$$

### Example 5.8: Removal of periodic noise by notch filtering



## Explanation of Example 5.8:



Each of the columns of the image above left contains a discretization  $f$  of the function  $f(x) = \sin(20x)$   $x \in [0, 2\pi]$ . The negative of the image is displayed.

The image in the middle shows the negative of `abs(fftshift(fft(f)))`

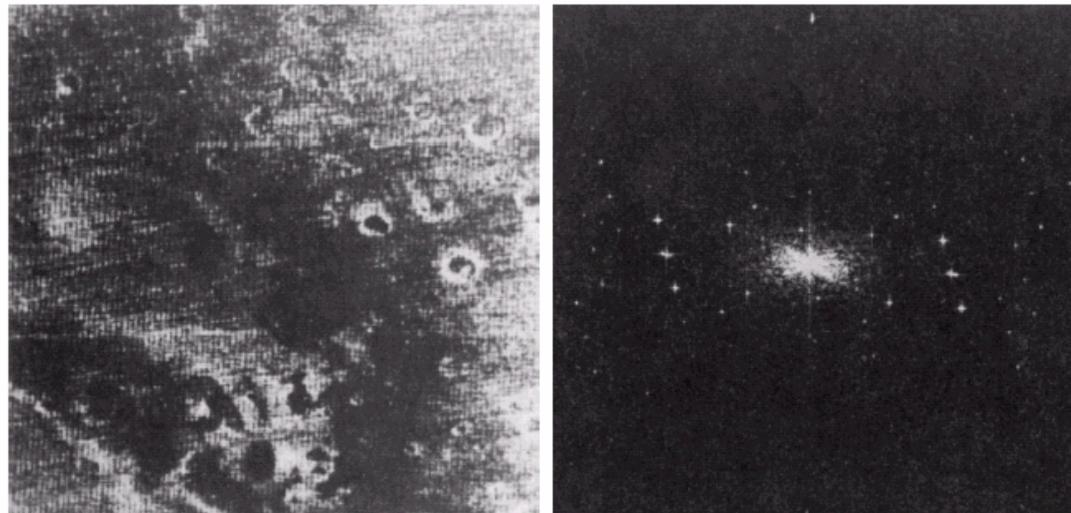
The image above right shows the negative of `abs(fftshift(fft2(f)))`

## 5.4.4 Optimum Notch Filtering

Starlike components in Fourier spectrum indicate more than one sinusoidal pattern

a b

**FIGURE 5.20**  
(a) Image of the Martian terrain taken by *Mariner 6*.  
(b) Fourier spectrum showing periodic interference.  
(Courtesy of NASA.)



Observe  $G(u, v)$  and experiment with different notch pass filters  $H_{\text{NP}}(u, v)$  where

$$\eta(x, y) = \text{IFT} \{ H_{\text{NP}}(u, v) G(u, v) \}$$

We want to optimize a weighting or modulation function  $w(x, y)$  where

$$\hat{f}(x, y) = g(x, y) - w(x, y) \eta(x, y) \quad (1)$$

in such a way that local variances of  $\hat{f}(x, y)$  is minimized



Consider a neighbourhood size of  $(2a + 1)$  by  $(2b + 1)$  about every point

Local variance of  $\hat{f}(x, y)$  at  $(x, y)$ :

$$\sigma^2(x, y) = \frac{1}{(2a + 1)(2b + 1)} \sum_{s=-a}^a \sum_{t=-b}^b [\hat{f}(x + s, y + t) - \overline{\hat{f}}(x, y)]^2 \quad (2)$$

$$\overline{\hat{f}}(x, y) = \frac{1}{(2a + 1)(2b + 1)} \sum_{s=-a}^a \sum_{t=-b}^b \hat{f}(x + s, y + t)$$

Substitute (1) into (2):

$$\sigma^2(x, y) = \frac{1}{(2a + 1)(2b + 1)} \sum_{s=-a}^a \sum_{t=-b}^b \left\{ [g(x + s, y + t) - w(x + s, y + t) \eta(x + s, y + t)] - [\overline{g}(x, y) - \overline{w(x, y) \eta(x, y)}] \right\}^2$$

Assume that  $w(x, y)$  remains constant over the neighbourhood, that is let

$$w(x + s, y + t) = w(x, y),$$

then

$$\overline{w(x, y) \eta(x, y)} = w(x, y) \overline{\eta(x, y)}$$

and

$$\sigma^2(x, y) = \frac{1}{(2a+1)(2b+1)} \sum_{s=-a}^a \sum_{t=-b}^b \{ [g(x+s, y+t) - w(x, y) \eta(x+s, y+t)] - [\bar{g}(x, y) - w(x, y) \bar{\eta}(x, y)] \}^2$$

To minimize  $\sigma^2(x, y)$ , we solve  $\frac{\partial \sigma^2(x, y)}{\partial w(x, y)} = 0$  for  $w(x, y)$ .

The result is

$$w(x, y) = \frac{\overline{g(x, y) \eta(x, y)} - \bar{g}(x, y) \bar{\eta}(x, y)}{\overline{\eta^2(x, y)} - [\bar{\eta}(x, y)]^2}$$

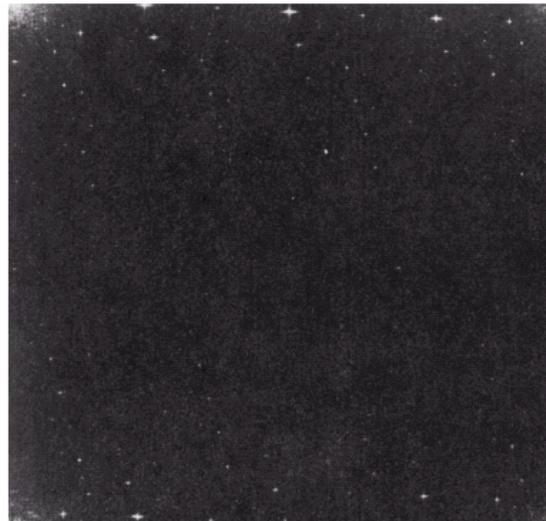
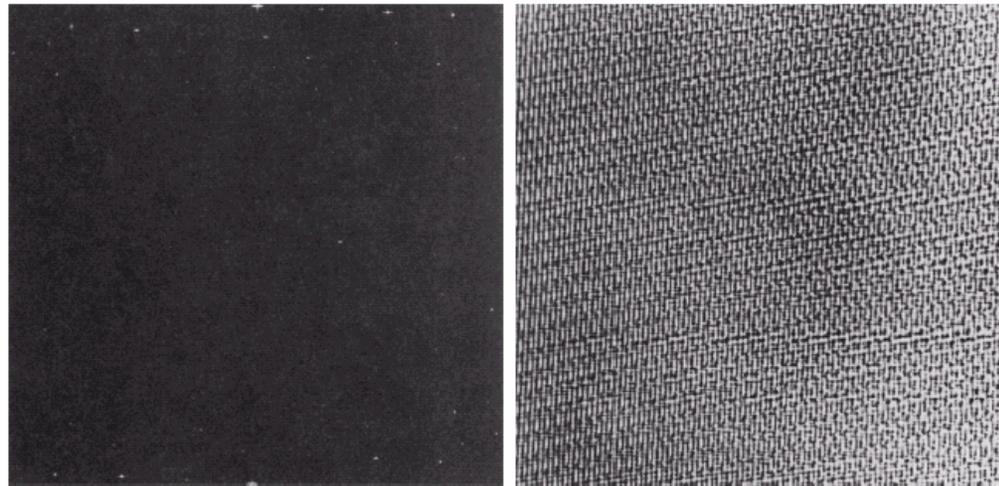
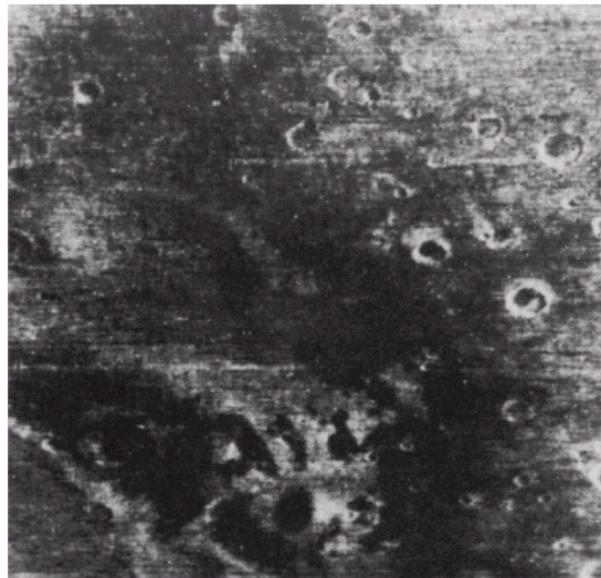


FIGURE 5.21 Fourier spectrum (without shifting) of the image shown in Fig. 5.20(a). (Courtesy of NASA.)



a b

**FIGURE 5.22** (a) Fourier spectrum of  $N(u, v)$ , and (b) corresponding noise interference pattern  $\eta(x, y)$ . (Courtesy of NASA.)



**FIGURE 5.23** Processed image. (Courtesy of NASA.)



## 5.5 Linear, position-invariant degradations

$$\begin{aligned}g(x, y) &= (h * f)(x, y) + \eta(x, y) \\ &= \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} f(\alpha, \beta) h(x - \alpha, y - \beta) d\alpha d\beta + \eta(x, y)\end{aligned}$$

$$G(u, v) = H(u, v) F(u, v) + N(u, v)$$

- Image deconvolution
- Deconvolution filters

## 5.6 Estimating the degradation function

- (1) Observation
- (2) Experimentation
- (3) Mathematical modelling

### 5.6.1 Estimation by image observation

- Given: only the degraded image
- Choose subimage that contains simple structures with little noise:  $g_s(x, y)$

- Estimate original subimage:  $\hat{f}_s(x, y)$

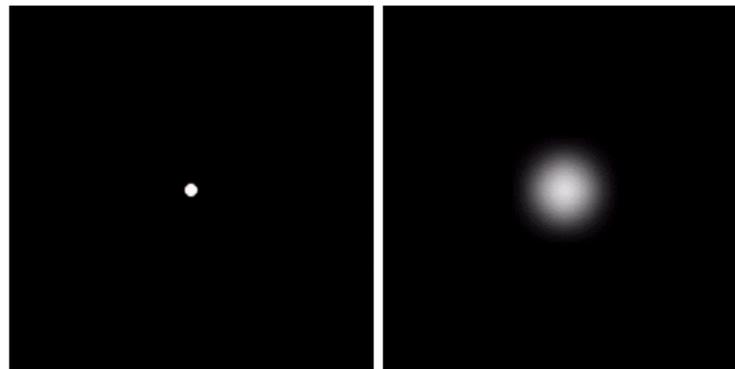
$$H_s(u, v) = \frac{G_s(u, v)}{\hat{F}_s(u, v)}$$

- Construct  $H(u, v)$  on a larger scale with similar “shape” as  $H_s(u, v)$

### 5.6.2 Estimation by experimentation

- Given: degraded image AND similar acquisition equipment
- Change settings until image resembles degraded image
- Acquire image of impulse (dot of light) with same settings

$$H(u, v) = \frac{G(u, v)}{A}, \text{ with } A \text{ the strength of impulse}$$



a b

**FIGURE 5.24**

Degradation estimation by impulse characterization. (a) An impulse of light (shown magnified). (b) Imaged (degraded) impulse.

## 5.6.3 Estimation by modelling

- Atmospheric turbulence

$$H(u, v) = e^{-k(u^2+v^2)^{5/6}}$$

a b  
c d

**FIGURE 5.25**

Illustration of the atmospheric turbulence model.

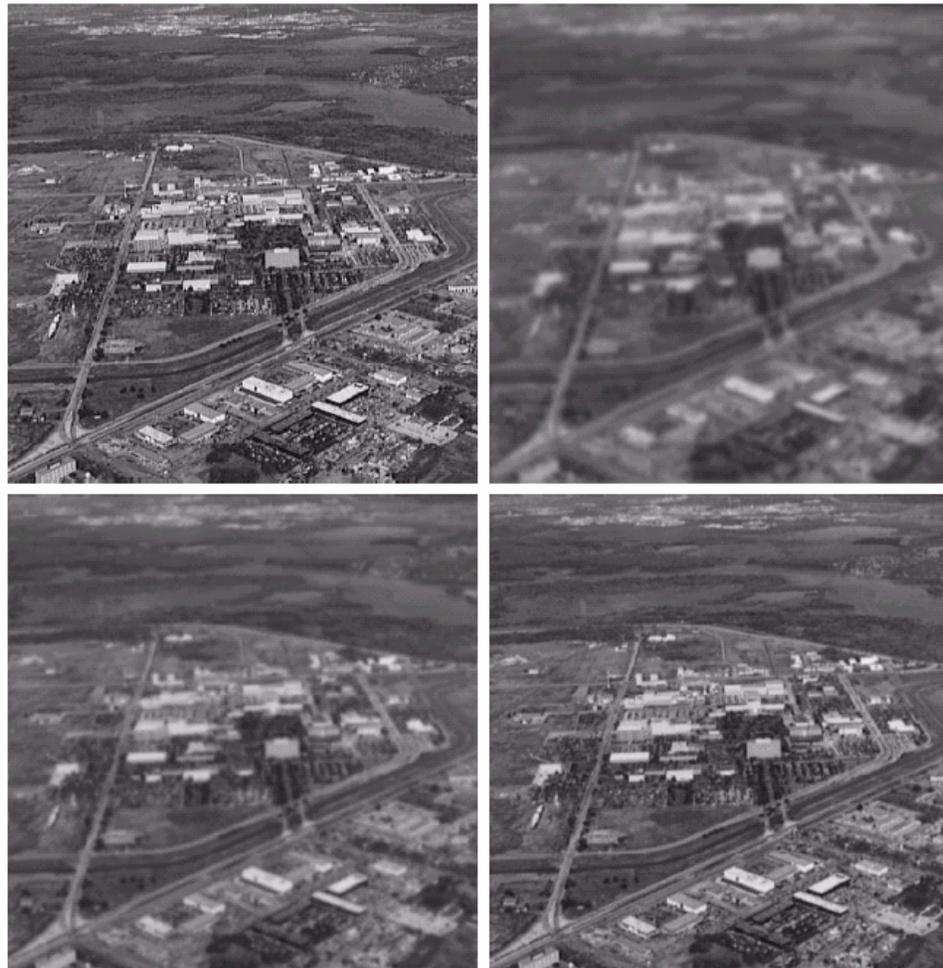
(a) Negligible turbulence.

(b) Severe turbulence,  $k = 0.0025$ .

(c) Mild turbulence,  $k = 0.001$ .

(d) Low turbulence,  $k = 0.00025$ .

(Original image courtesy of NASA.)





- **Basic mathematical principles: uniform linear motion**

Let  $T$  be duration of exposure, then  $g(x, y) = \int_0^T f[x - x_0(t), y - y_0(t)] dt$

$$\begin{aligned} G(u, v) &= \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} g(x, y) e^{-2\pi i(ux+vy)} dx dy \\ &= \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} \left( \int_0^T f[x - x_0(t), y - y_0(t)] dt \right) e^{-2\pi i(ux+vy)} dx dy \\ &= \int_0^T \left( \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} f[x - x_0(t), y - y_0(t)] e^{-2\pi i(ux+vy)} dx dy \right) dt \\ &= \int_0^T F(u, v) e^{-2\pi i[ux_0(t)+vy_0(t)]} dt \\ &= F(u, v) \int_0^T e^{-2\pi i[ux_0(t)+vy_0(t)]} dt \end{aligned}$$

**Define**  $H(u, v) = \int_0^T e^{-2\pi i[ux_0(t)+vy_0(t)]} dt$  **so that**  $G(u, v) = H(u, v) F(u, v)$

**If**  $x_0(t)$  **and**  $y_0(t)$  **are known, then**  $H(u, v)$  **is known**



## Illustration

Let  $x_0(t) = \frac{at}{T}$  and  $y_0(t) = 0$

Note that  $x_0(T) = a$

$$\begin{aligned} H(u, v) &= \int_0^T e^{-2\pi i u x_0(t)} dt \\ &= \int_0^T e^{-2\pi i u at/T} dt \\ &= \vdots \\ &= \frac{T}{\pi u a} \sin(\pi u a) e^{-\pi i u a} \end{aligned}$$

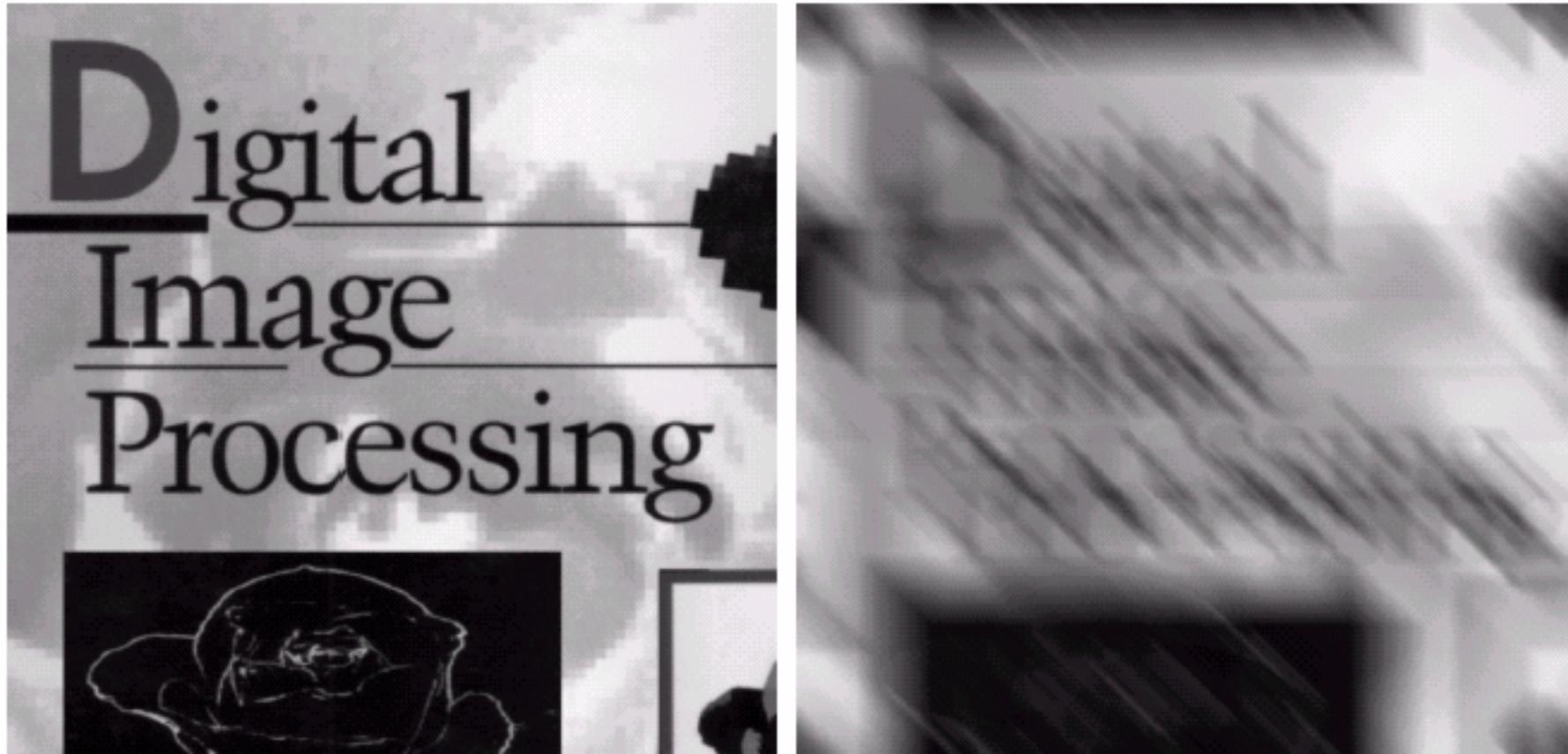
(verify)

Note that  $H(u, v) = 0$  if  $u = \frac{n}{a}$ ,  $n \in \mathbb{Z}$

When  $x_0(t) = \frac{at}{T}$  and  $y_0(t) = \frac{bt}{T}$ , then

$$H(u, v) = \frac{T}{\pi(ua + vb)} \sin[\pi(ua + vb)] e^{-\pi i(ua + vb)}$$

**Example 5.10:** Image blurring due to motion



a b

**FIGURE 5.26** (a) Original image. (b) Result of blurring using the function in Eq. (5.6-11) with  $a = b = 0.1$  and  $T = 1$ .